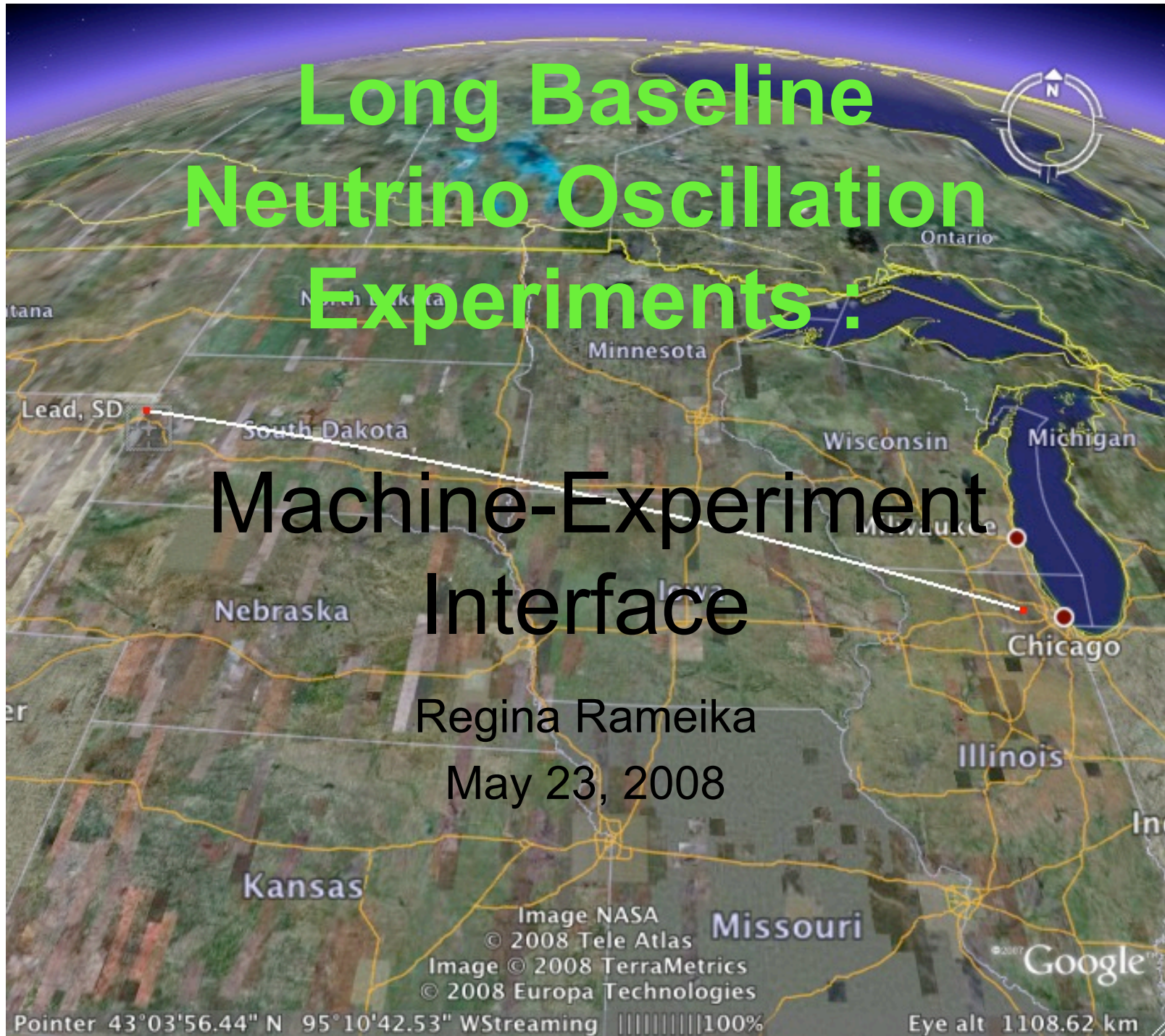


Long Baseline Neutrino Oscillation Experiments :

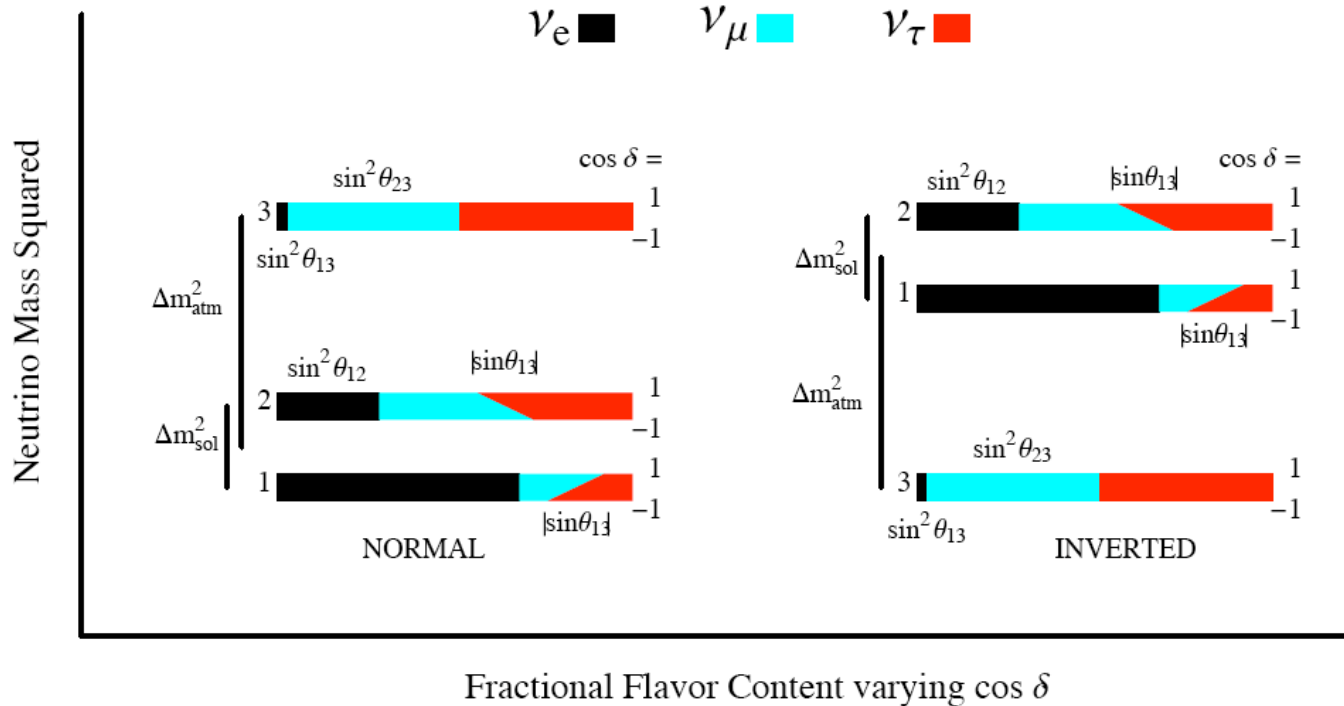
Machine-Experiment Interface

Regina Rameika

May 23, 2008



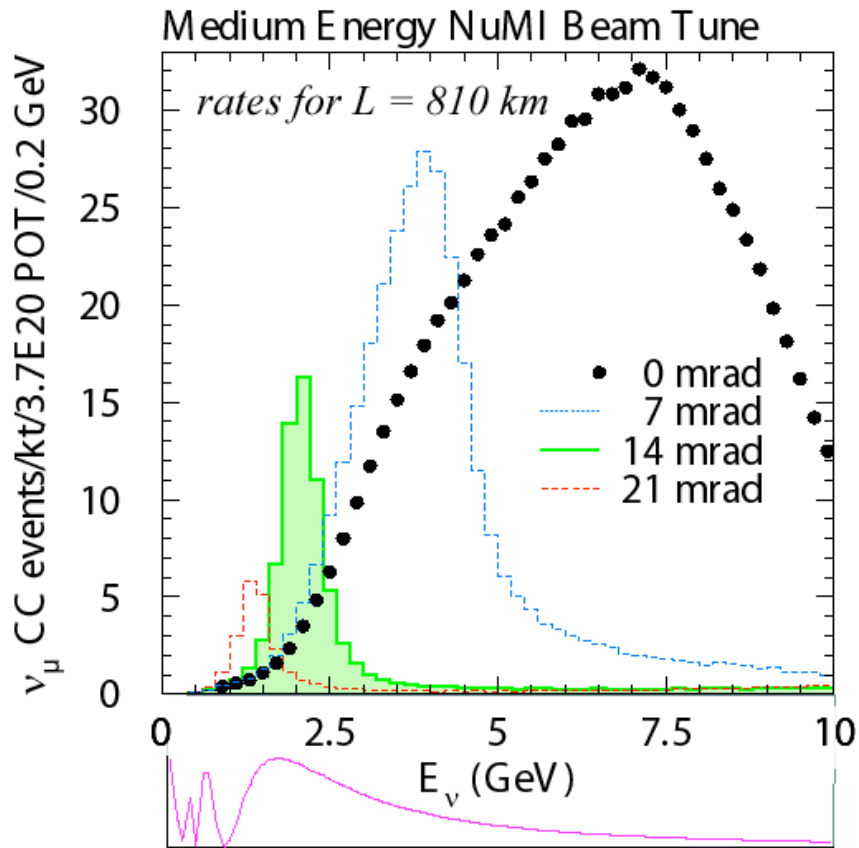
The Questions for Long Baseline Neutrino Oscillation Experiments



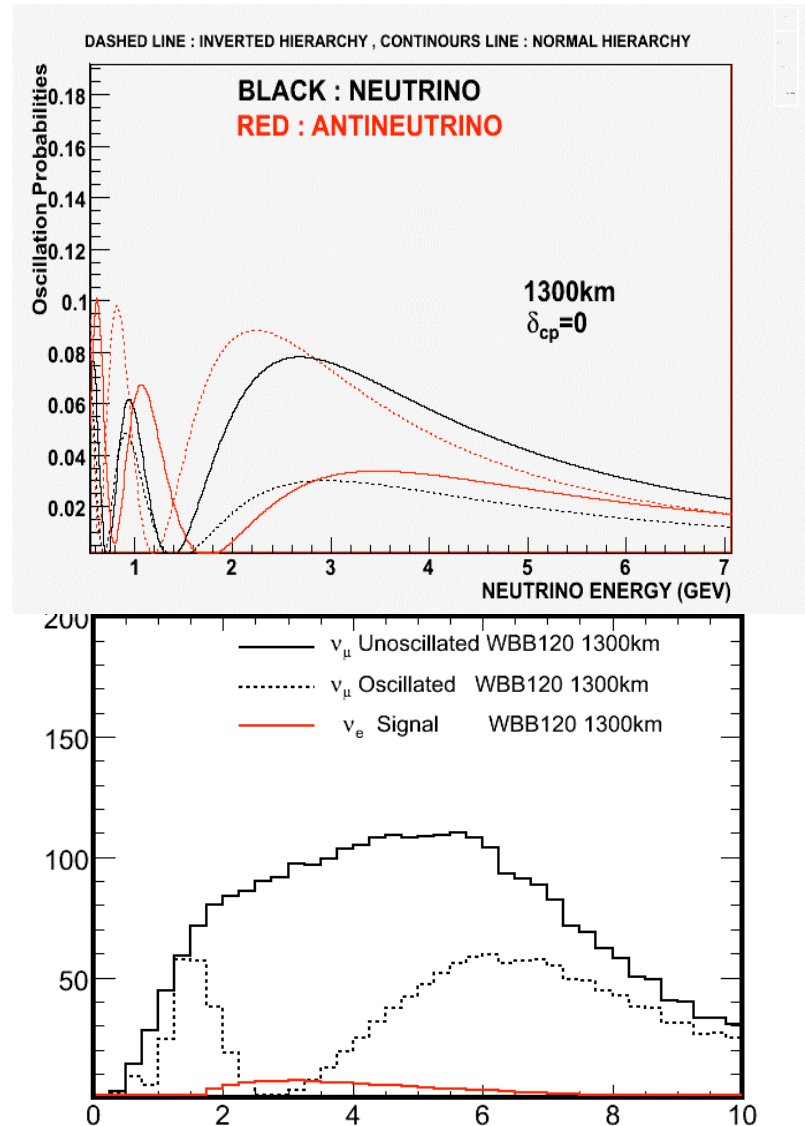
- What is the size of θ_{13} ?
- Is the hierarch normal or inverted?
- What is the value of δ_{CP} ?

Experimental Techniques

Off-axis/narrow band beam

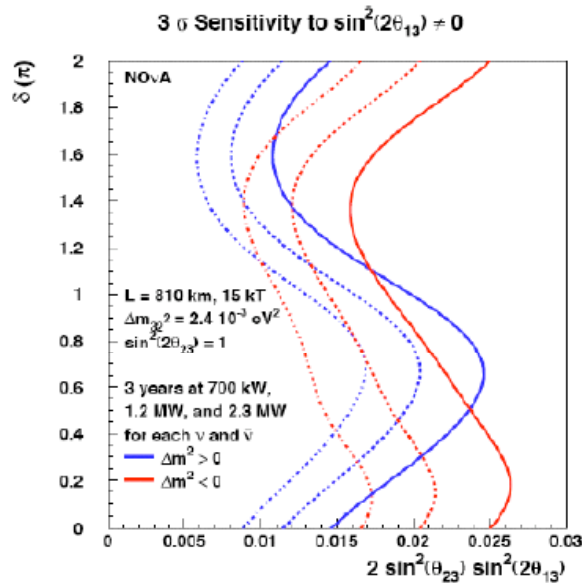


On-axis/wide band beam

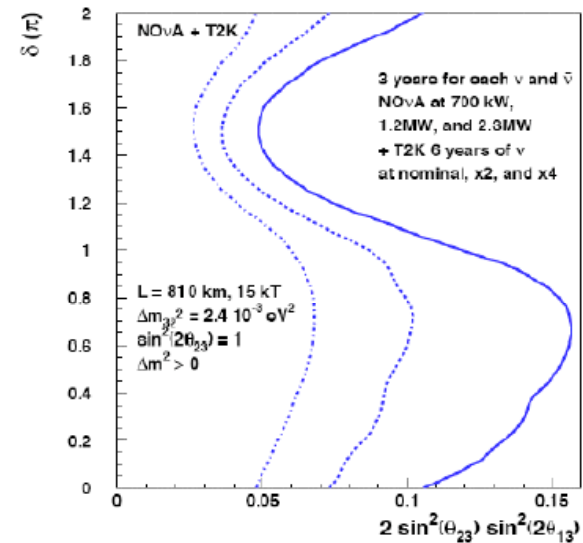


An experiment's sensitivity is measured in $\sin^2 2\theta_{13}$ and δ_{CP} space

NOvA Experiment Sensitivities (shown by G.Feldman at P5 2008 SLAC meeting)



95% C.L. resolution of the mass hierarchy



Normal Ordering

Lines ~represent equal numbers

Of signal events for a given set of $\sin^2 2\theta_{13}, \delta_{CP}$

N. Saoulidou, Fermilab, February 02
 2008

More neutrinos or more mass;
 more neutrinos \rightarrow instantaneous intensity+ time

		Neutrino Rates				Anti Neutrino Rates			
Beam (mass ordering)	$\sin^2 2\theta_{13}$	δ_{CP} deg.							
		0°	-90°	180°	+90°	0°	-90°	180°	+90°
NuMI LE 12 km offaxs (+)	0.02	76	108	69	36	20	7.7	17	30
NuMI LE 12 km offaxs (-)	0.02	46	77	52	21	28	14	28	42
NuMI LE 12 km offaxs (+)	0.1	336	408	320	248	86	57	78	106
NuMI LE 12 km offaxs (-)	0.1	210	280	224	153	125	95	126	157
NuMI LE 40 km offaxs (+)	0.02	5.7	8.8	5.1	2.2	2.5	1.6	0.7	3.3
NuMI LE 40 km offaxs (-)	0.02	4.2	8.0	5.7	2.0	2.3	2.2	0.8	3.6
NuMI LE 40 km offaxs (+)	0.1	17	24	15	9.4	6.7	2.8	4.6	8.5
NuMI LE 40 km offaxs (-)	0.1	12	21	16	7.7	6.6	3.4	6.4	9.6
WBLE 1300 km (+)	0.02	141	192	128	77	19	11	18	36
WBLE 1300 km (-)	0.02	58	111	88	35	45	25	45	64
WBLE 1300 km (+)	0.1	607	720	579	467	106	67	83	122
WBLE 1300 km (-)	0.1	269	388	335	216	196	154	196	240
WBLE 2500 km (+)	0.02	61	103	88	46	11	4.6	4.7	11
WBLE 2500 km (-)	0.02	16	36	33	13	28	15	18	31
WBLE 2500 km (+)	0.1	270	361	328	238	27	13	13	28
WBLE 2500 km (-)	0.1	47	92	85	39	103	74	80	109

Charge current
events per
100kT mass
per 1 MW per
10⁷ sec

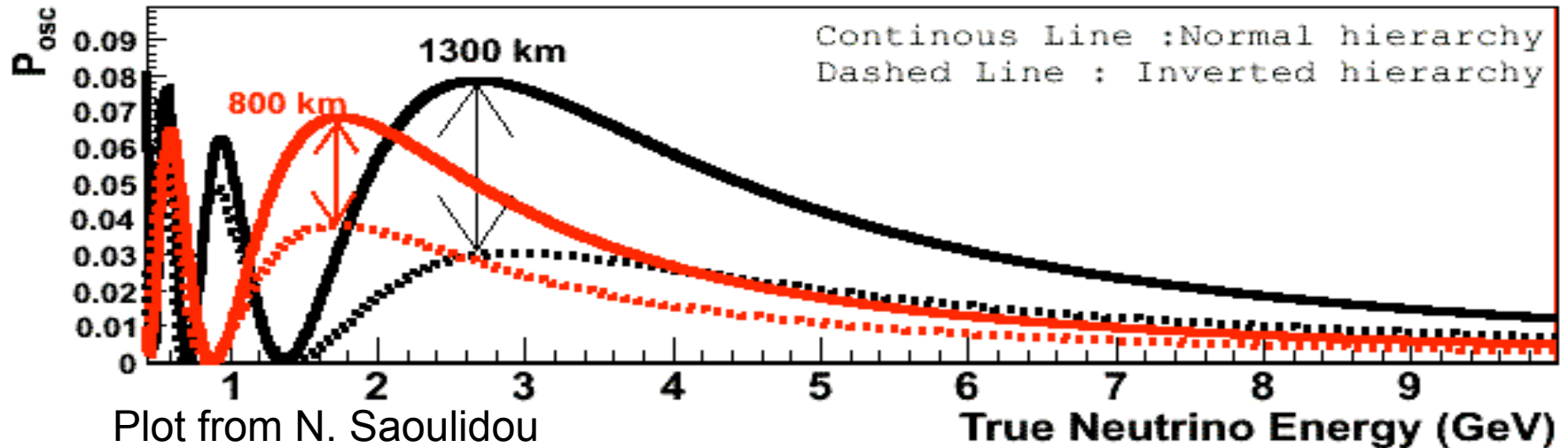
No detector model
or backgrounds

(NuMI - 120 GeV
WBLE - 60 GeV)

DUSEL
rates

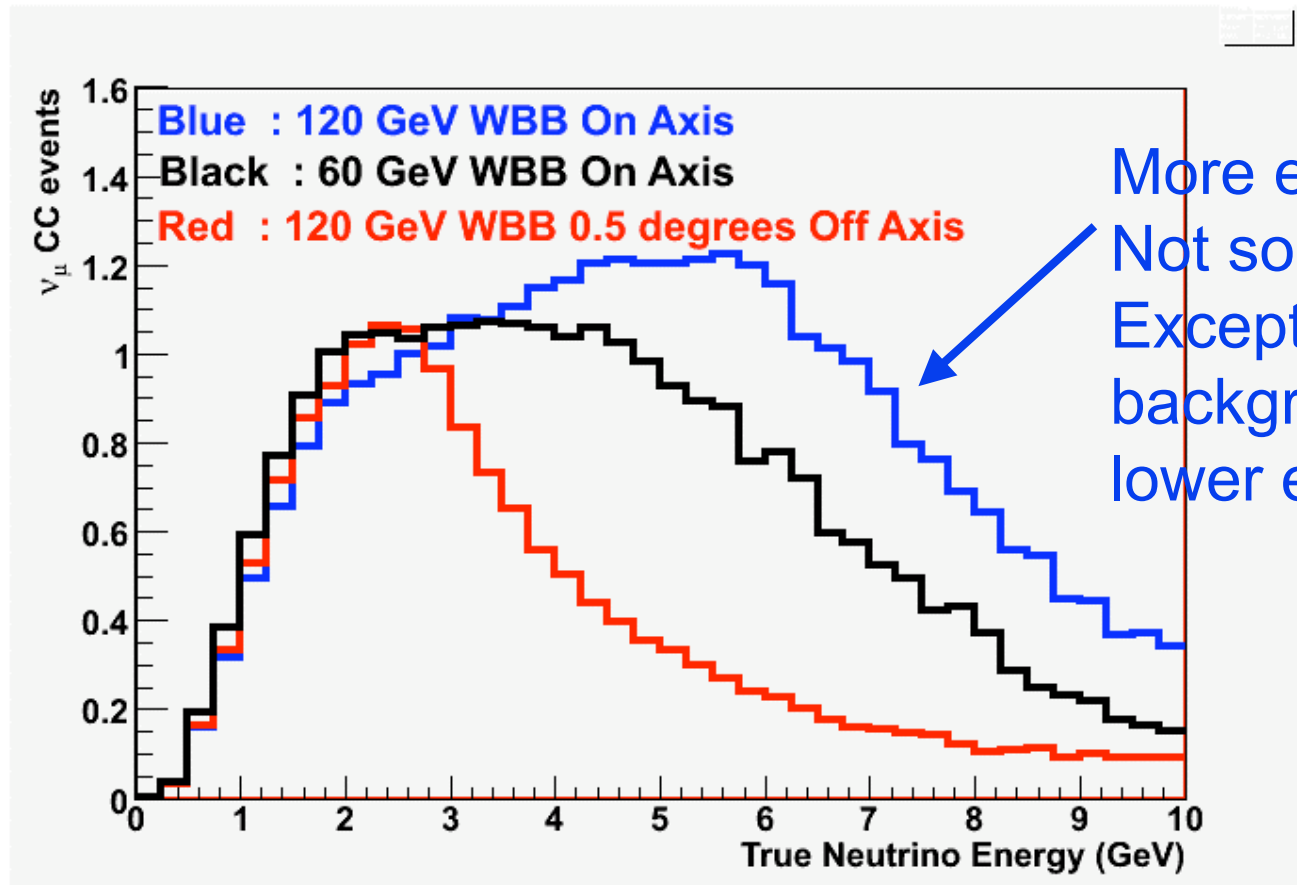
~10-1000 evts

Why is the longer baseline so much better?

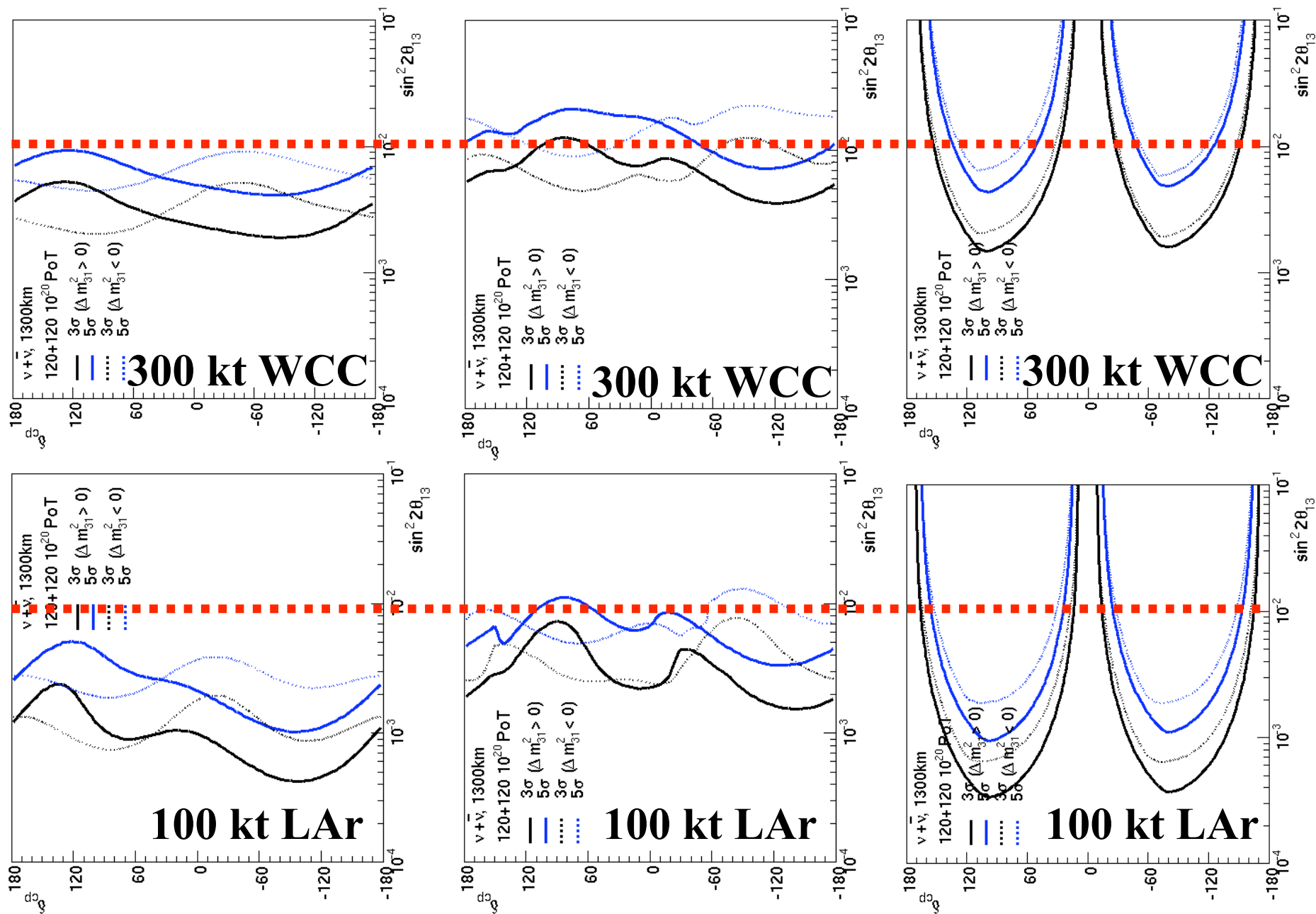


- Oscillation maxima are moved to higher energy
- Matter effects are significantly larger

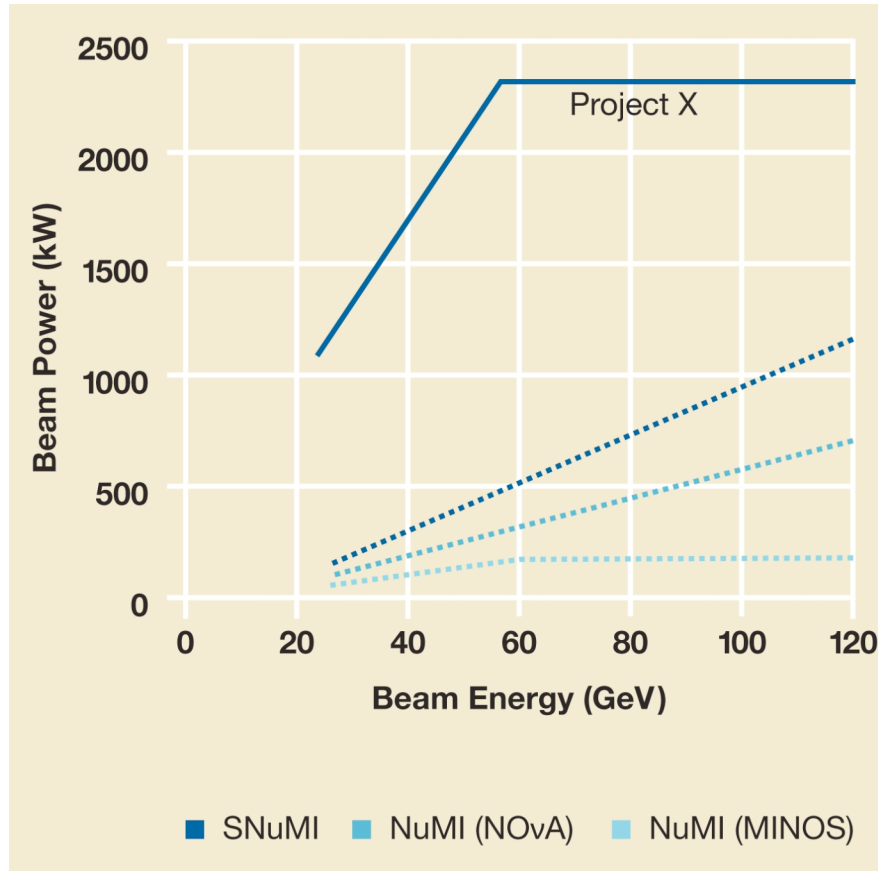
Options and optimizations



100 kt LAr vs 300 kt of WCC



- 60 -120 GeV protons from the Main Injector fed by Project X



20-40x10²⁰ POT/yr

10x10²⁰ POT/yr

6x10²⁰ POT/yr

3x10²⁰ POT/yr

Recent sensitivity studies are being done for 120x10²⁰ POT each ν and $\bar{\nu}$ (120 GeV)

$$POT(10^{20}) = \frac{1000 \times BeamPower(MW) \times T(10^7 s)}{1.602 \times E_p(GeV)}$$

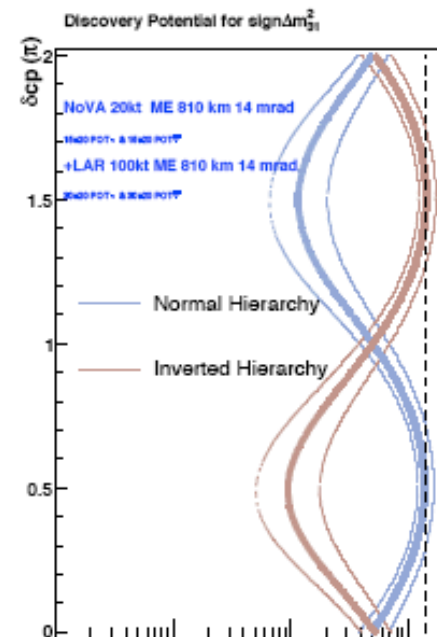
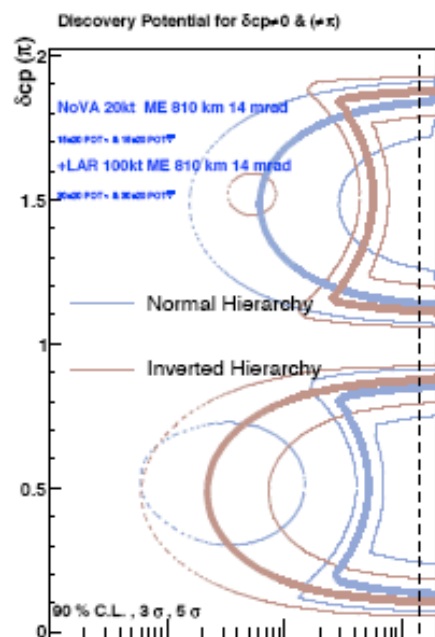
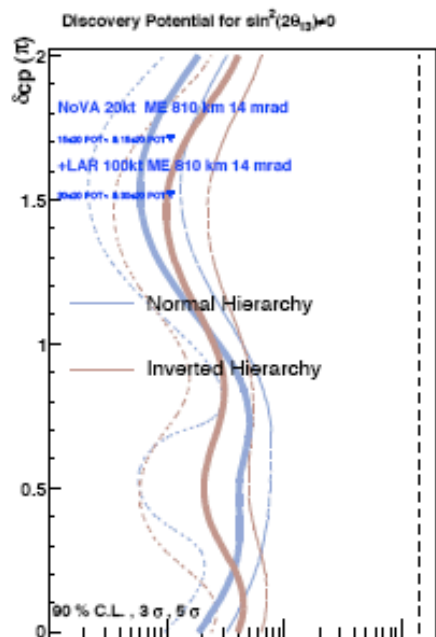
Summary

- Flexibility to choose energy at the same power is a real plus
 - Lower proton energy may have real advantages
- Integrated intensity goal
 - 240 (480) $\times 10^{20}$ POT @ 120 (60) GeV
- Annual intensity goal
 - 20 (40) $\times 10^{20}$ POT @ 120 (60) GeV
- Instantaneous intensity
 - Single turn extraction (10 μ s every 1.4 sec)
 - No rate effects at far detector
 - 15×10^{14} per spill \rightarrow 2 MW at 120 GeV

Summary

- Experiments are statistics limited for values of $\sin^2 2\theta_{13}$ well below 0.01
- For far detector operation, fast extraction ($10\mu\text{s}$ every 1.4 sec) of 15×10^{14} per spill yields $\sim 40 <$ interactions per year < 4000
- Sensitivities achievable at the statistics/systematics limit can use up to 120×10^{20} protons in each the neutrino and anti-neutrino running mode \rightarrow 5 years each mode
- For near detector operation at this intensity, event overlap could be significant; detector configuration will need to take this into account

NuMI 1st max



NuMI 2nd max

